# Pearson Edexcel 

Mark Scheme (Results)

Summer 2019

Pearson Edexcel GCSE In Astronomy (1AS0) Paper 1: Naked eye Astronomy

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Summer 2019
Publications Code 1ASO_01_1906_MS
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- All candidates must receive the same treatment. Examiners must mark the first candidate in exactly the same way as they mark the last.
- Mark schemes should be applied positively. Candidates must be rewarded for what they have shown they can do rather than penalised for omissions.
- Examiners should mark according to the mark scheme not according to their perception of where the grade boundaries may lie.
- There is no ceiling on achievement. All marks on the mark scheme should be used appropriately.
- All the marks on the mark scheme are designed to be awarded. Examiners should always award full marks if deserved, i.e. if the answer matches the mark scheme. Examiners should also be prepared to award zero marks if the candidate's response is not worthy of credit according to the mark scheme.
- Where some judgement is required, mark schemes will provide the principles by which marks will be awarded and exemplification may be limited.
- When examiners are in doubt regarding the application of the mark scheme to a candidate's response, the team leader must be consulted.
- Crossed out work should be marked UNLESS the candidate has replaced it with an alternative response.

\begin{tabular}{|c|c|c|}
\hline Question number \& Answer \& Mark \\
\hline 1(a) \& \begin{tabular}{l}
(i) A Cassiopeia \\
\(B\) is incorrect because the stars are not in the shape of the constellation Cassiopeia. \\
C is incorrect because the stars are not in the shape of the constellation Cassiopeia. \\
\(D\) is incorrect because the stars are not in the shape of the constellation Cassiopeia. \\
(ii) C Pegasus \\
A is incorrect because the stars are not in the shape of the constellation Pegasus. \\
\(B\) is incorrect because the stars are not in the shape of the constellation Pegasus. \\
D is incorrect because the stars are not in the shape of the constellation Pegasus. \\
(iii) B Cygnus \\
A is incorrect because the stars are not in the shape of the constellation Cygnus. \\
C is incorrect because the stars are not in the shape of the constellation Cygnus. \\
\(D\) is incorrect because the stars are not in the shape of the constellation Cygnus.
\end{tabular} \& (1)

(1)

(1) <br>
\hline
\end{tabular}

| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 1(b) | (i) C Meteor <br> A is incorrect because it lasts for much more than a few seconds. <br> $B$ is incorrect because it lasts for much more than a few seconds. <br> $D$ is incorrect because it lasts for much more than a few seconds. <br> (ii) B Comet <br> A is incorrect because it is not star-like with a fuzzy tail. <br> C is incorrect because it is not star-like with a fuzzy tail. <br> D is incorrect because it is not star-like with a fuzzy tail. | (1) <br> (1) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 1(c) | Outline of solar disc during partial eclipse, as in example below. <br> (Black background obviously not required). | (1) |
|  | Accept any orientation. <br> Accept any size of solar disc. <br> Line of intersection must be curved. <br> Reject: <br> Annular eclipse. |  |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 2(a) | (i) C Epicycles <br> A is incorrect because it is an observation. <br> $B$ is incorrect because it was proposed by Kepler <br> $D$ is incorrect because it is an ancient idea. <br> (ii) B Elliptical orbits <br> A is incorrect because it is an observation. <br> C is incorrect because it was proposed by Ptolemy <br> $D$ is incorrect because it is an ancient idea. | (1) <br> (1) |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 2(b) | (i) C Always be above the horizon <br> A is incorrect because this is not the behaviour of a circumpolar object. <br> $B$ is incorrect because this is not the behaviour of a circumpolar object. <br> $D$ is incorrect because this is not the behaviour of a circumpolar object. <br> (ii) C $21^{\text {st }}$ September <br> A is incorrect because it is the date of the Spring Equinox <br> B is incorrect because it is the date of the Summer Solstice <br> D is incorrect because it is the date of the Winter Solstice | (1) <br> (1) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 2(c) | (i) $\quad$B occultation <br> A is incorrect because it does not describe an <br> occultation. <br> C is incorrect because it does not describe an <br> occultation. <br> D is incorrect because it does not describe an <br> occultation. <br> (ii) $\mathbf{C}$ opposition <br> A is incorrect because it does not describe an <br> occultation. <br> B is incorrect because it does not describe an <br> occultation. <br> D is incorrect because it does not describe an <br> occultation. | (1) |


| Question <br> number | Answer | Mark |
| :--- | :---: | :--- |
| 2(d) | (i)A double <br> B is incorrect because the force of gravity is <br> proportional to mass <br> C is incorrect because the force of gravity is <br> proportional to mass <br> Dis incorrect because the force of gravity is <br> proportional to mass <br> (ii)D drop to one quarter <br> A is incorrect because the force of gravity is <br> inversely proportional to distance squared <br> B is incorrect because the force of gravity is <br> inversely proportional to distance squared <br> C is incorrect because the force of gravity is <br> inversely proportional to distance squared$\quad$(1) |  |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 3(a) | S <br> (i) S (ignore any arrows) <br> (ii) O <br> (iii) P (ignore arrow length; reject any points) | (1) <br> (1) <br> (1) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 3(b) | (i) Not a likely explanation <br> Ancient monuments were generally closely aligned <br> astronomically | (1) <br> (1) |
|  | (ii) Not a likely explanation <br> All visible stars formed long before Earth. | (1) |
|  | (iii) Not a likely explanation <br> Precession would move all three stars together, not relative <br> to each other | (1) |
|  |  | (1) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 4(a) | Summer/Mid-summer |  |
|  | Solstice | $\mathbf{( 1 )}$ |
|  | 'Midsummer's Day' therefore scores 1. | $\mathbf{( 1 )}$ |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 4(b) | Rises and sets / Travels East to West <br> Through Alice's overhead point/zenith <br> Insufficient: 'highest point' | $\mathbf{( 1 )}$ |
| $\mathbf{( 1 )}$ |  |  |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 4(c) | Point 'S'/this declination is overhead <br> (from) Tropic of Cancer | $\mathbf{1 1 )}$ <br> $\mathbf{( 1 )}$ |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 4(d) | Precession <br> moves the stars relative to/'behind' point 'S' (on ecliptic). | $\mathbf{( 1 )}$ <br> $\mathbf{( 1 )}$ |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 4(e) | On ecliptic OR: <br> Sun/Moon/planets pass through them | $\mathbf{( 1 )}$ |


| Question |
| :--- | :--- |
| number | Answer


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 5(b) | $3300000(\mathrm{~km})$ Allow: 3325000 <br> $(x 19=1)$ | (2) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 5(c) | $98 \%(97.8 \%)$ <br> $3325000-150$ million $=146675000$ | (2) <br> $(1)$ |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 5(d) | Measuring angle close to $90^{\circ}$. <br> Compound error from using earlier measurements. | $\mathbf{( 1 )}$ |
| $\mathbf{( 1 )}$ |  |  |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| $\mathbf{6 ( a ) \mathbf { i }}$ | Star reaches highest point/crosses meridian/is due south <br> Labelled diagram (required) shows this. | $\mathbf{( 1 )}$ <br> $\mathbf{( 1 )}$ |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 6(a) ii | Curved (non-flat) surface of Earth. <br> Allow: different latitudes | (1) |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 6(a) iii | 28 800km <br> Any 1: <br> - Angular difference $=712^{\circ}-0^{\circ}=712^{\circ}$ <br> - $360^{\circ} / 712_{2}{ }^{\circ}=48 x$ <br> - $48 \times 600 \mathrm{~km}$ | (2) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 6(a) iv | Any 2 from: <br> - Observe from further south to increase angle <br> (Reject: observe star (Canopus) from further north) <br> - Choose another star to avoid observing on horizon <br> Obtain a more accurate value for Rhodes-Alexandria <br> distance <br> Repeat and average observations / use multiple stars | (2) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| $\mathbf{6 ( b )}$ | $\mathbf{5 8 0}(577.8)(\mathbf{k m})$ | $\mathbf{( 2 )}$ |
|  | Any 1 from: <br> $360 / 5.2(=69(69.23))$ <br>  <br>  <br> $40000 / 69(=580)$ | $(1)$ |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 7(a) | (i) D 11:33 <br> A is incorrect because it is the Apparent Solar Time $B$ is incorrect because it is the Local Mean Time $C$ is incorrect because it is not related to the question <br> (ii) D 11:33 <br> A is incorrect because it is the Apparent Solar Time $B$ is incorrect because it is the Local Mean Time C is incorrect because it is not related to the question <br> (iii) B 11:21 <br> A is incorrect because it is the Apparent Solar Time $C$ is incorrect because it is not related to the question <br> D is incorrect because it is the Greenwich Mean Time <br> (iv) A 11:15 <br> $B$ is incorrect because it is the Local Mean Time $C$ is incorrect because it is not related to the question D is incorrect because it is the Greenwich Mean Time | (1) <br> (1) <br> (1) <br> (1) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 7(b) | Identification one of the following causes. <br> Explanation of how it introduces a difference between Solar and <br> Mean time. <br> Eccentricity of Earth's Orbit. Changing distance between Earth <br> and Sun means changes in Earth's orbital speed resulting in <br> changes in the Sun's apparent speed across the sky. | $\mathbf{( 1 )}$ <br> (1) <br> Obliquity of Ecliptic. Tilt between Earth's orbital plane and spin <br> axis changes apparent altitude of Sun in sky and thus its angular <br> speed across the sky. |


| Question number | Indicative content | Mark |
| :---: | :---: | :---: |
| 7(c) | - Equation of Time is difference between Apparent Solar Time and Mean Solar Time <br> - Apparent Solar Time can be found from a sundial or shadow stick <br> - Mean Solar Time can be found from a clock <br> - Effect of longitude on AST needs to be allowed for <br> - Need to ensure no direct viewing of Sun <br> - Easier to use MST as independent variable as it is easier to assess when it has reached a whole number of minutes (unlike AST). | (6) |


| Level | Mark | Descriptor |
| :---: | :---: | :---: |
|  | 0 | No rewardable material. |
| Level 1 | 1-2 | - Presents a simple observing programme showing some awareness of the phenomenon under investigation <br> - Identifies some feasible parameters such as location, and observing times. <br> - No evidence of awareness of need for repeated observations or observation over an extended time period to achieve reliable data. |
| Level 2 | 3-4 | - Presents a sound observing programme showing clear awareness of the phenomenon under investigation <br> - Identifies a number of feasible parameters such as location, and observing times. <br> - Shows awareness of need for repeated observations or observation over an extended time period to achieve reliable data. |
| Level 3 | 5-6 | - Presents a detailed observing programme showing a thorough understanding of the phenomenon under investigation <br> - Identifies all relevant parameters such as location, and observing times. <br> - Shows a clear understanding of the need for repeated observations or observation over an extended time period to achieve reliable data and reflects this clearly in their design. |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 8(a) | • Appropriate and correctly-labelled scales on both axes | $\mathbf{( 1 )}$ |
|  | • Points correctly plotted. | $\mathbf{( 1 )}$ |
|  | • Smooth U-shaped curve through points. | $\mathbf{( 1 )}$ |


| Question number | Indicative content | Mark |
| :---: | :---: | :---: |
| 8(b) | - Readings taken either side of local noon <br> - Measurements taken on whole numbers of minutes <br> - Intervals in readings are too large <br> - Around noon, the Sun is moving one or two degrees between readings <br> - An error of even one degree in latitude represents a substantial distance at sea [ $\sim 70$ miles] <br> - Using the Sun's altitude as the independent variable may have been a more effective method <br> - Altitude of Sun at noon correctly assessed from data ( $42^{\circ}$ ) <br> - Latitude calculation is incorrect: Co-latitude $+8^{\circ}=42^{\circ}$, giving correct latitude of $56^{\circ}$. | (6) |


| Level | Mark | Descriptor |
| :---: | :---: | :---: |
|  | 0 | No rewardable material. |
| Level 1 | 1-2 | - A few inadequacies in the data are noted <br> - A few shortcomings of the method used are identified <br> - Some mention of relevant astronomical theory is made <br> - At least one feasible suggestion for improving the method is made. |
| Level 2 | 3-4 | - The major inadequacies in the data are noted <br> - These are each linked to a particular shortcoming of the method used are identified <br> - Relevant astronomical theory is used <br> - Feasible suggestions for improving the method are made. |
| Level 3 | 5-6 | - All inadequacies in the data are noted <br> - These are each linked to a particular shortcoming of the method used are identified <br> - Relevant astronomical theory is used to justify each of the above points <br> - Detailed suggestions for improving the method are made by systematically addressing each of the identified issues. |


| Question number | Answer | Additional guidance | Mark |
| :---: | :---: | :---: | :---: |
| 8(c) | (i) $20^{\circ}$ <br> Any 1 from: $\begin{aligned} & 13: 20-12: 00=1 \mathrm{~h} 20 \mathrm{~m} \\ & 1 \mathrm{~h} 20 \mathrm{~m} \times 15\left(=20^{\circ}\right) \end{aligned}$ <br> (ii) $150^{\circ}$ <br> W | (ignore any indication of +/- or direction of longitude difference) <br> (allow $+150^{\circ}$ ) <br> Allow ECF from (i) | (2) <br> (1) <br> (1) |


| Question <br> number | Answer | Additional guidance | Mark |
| :--- | :--- | :--- | :--- |
| 8(d) | (John) Harrison <br> (Ignore first name) | Allow: Harrison, <br> Harrisson or Harisson. | (1) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 9(a) | Any $\mathbf{2}$ of the following points, established by diagram or otherwise: <br> - RA of observer's meridian is 10 h 42 m <br> - First Point of Aries is 10 h 42 m from meridian <br> - Aldebaran is 6 h 6 m from meridian <br> - RA of Aldebaran $=10 \mathrm{~h} 42 \mathrm{~m}-6 \mathrm{~h} 6 \mathrm{~m}$ | (2) |
|  | - i.e. 4 h 36 m from First Point of Aries. |  |
|  | Diagram supporting on of the above. |  |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 9(b) | Any $\mathbf{1}$ of: <br> $\bullet$ Altitude of Celestial Equator on meridian is $90^{\circ}-42^{\circ}=48^{\circ}$ <br> $\bullet$ Altitude of Aldebaran $\left(64^{\circ} 30^{\prime}\right)$ therefore $=\operatorname{Dec}+48^{\circ}$ | (1) |
|  | Establishing NCP as $42^{\circ}$ above northern horizon and <br> Aldebaran's co-declination as $138^{\circ}-64^{\circ} 30^{\prime}=73^{\circ} 30^{\prime}$. |  |
|  | Labelled diagram illustrating one of the above. | (1) |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 9(c) | $54^{\circ} 45^{\prime}$ <br> Up to 1 'working' marks available for incorrect answers: <br> - Cel Eq is $38^{\circ} 15^{\prime}$ above southern horizon in Oxford <br> - Aldebaran's declination will put it a further $16^{\circ} 30^{\prime}$ above the horizon. | (2) |


| Question <br> number | Answer | Mark |
| :--- | :--- | :--- |
| 9(d) | (i) $\mathbf{4 h} \mathbf{3 6 m}$ <br> LST $=$ RA of objects on meridian, i.e. $10 \mathrm{~h} 42 \mathrm{~m}-6 \mathrm{~h} 6 \mathrm{~m}$ | (2) <br> (1) |
|  | (ii) 4 h 36 m <br> Mention of LST or RA of objects on meridian | (2) <br> (1) |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 10(a)(i) | Points established by diagram or otherwise: <br> - Solar eclipse occurs when Moon is between Earth and Sun <br> - Lunar eclipse occurs when Moon is on opposite side of Earth to Sun <br> - They can therefore be separated by half of a lunar orbit = $1 / 2 \times 28$ days. | (1) <br> (1) <br> (1) |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 10(a)(ii) | Points established by diagram or otherwise: <br> - Moon is much smaller than cone of shadow behind Earth in a lunar eclipse <br> - Moon fits almost exactly into line of sight 'cone' of rays to Sun in a solar eclipse <br> - Comparison of shadow areas on Earth's surface (night hemisphere of Earth v. tip of shadow cone). | (1) <br> (1) <br> (1) |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 10(b)(i) | 23 (22.9) days <br> Up to 2 'working' marks available for incorrect answers: <br> - Attempt to calculate $T^{2} / r^{3}$ for Moon's current orbit <br> - Correct value for $T^{2} / r^{3}$, e.g. $0.00373,745.29$ etc <br> - Correct substitution for new orbit, giving $T^{2} / 52^{3}=$ their value. <br> Correct Method: <br> Current lunar orbit: <br> $\mathrm{r}=380000 \mathrm{~km} / 6500 \mathrm{~km}=58.46$ Earth radii. <br> $\mathrm{T}^{2} / \mathrm{r}^{3}=(27.3)^{2} /(58.46)^{3}=\underline{0.00373}$ <br> (or equivalent in other units, such as 745.29 days $^{2} /$ Moon orbital radii ${ }^{3}$ ) <br> Lunar orbit of 52 Earth radii: <br> $r=52$, giving; $\begin{aligned} & T^{2} / 52^{3}=0.00373 \\ & T=\sqrt{ }\left(0.00373 \times 52^{3}\right) \\ & T=22.9 \text { days } \end{aligned}$ | (3) |


| Question number | Answer | Mark |
| :---: | :---: | :---: |
| 10(b)(ii) | 580000 (581 100) km <br> Up to 3 'working' marks available for incorrect answers: <br> - $T^{2} / r^{3}=0.00373$ (or their value from (i)) <br> - Doubling Moon's sidereal period ( $2 \times 27.3=54.6$ days) <br> - Calculating r=90.4 Earth radii <br> - Subtracting one Earth radius to find distance from Earth's surface. <br> Correct solution: <br> $\mathrm{T}^{2} / \mathrm{r}^{3}=0.00373$ (or their value from (i)) $\begin{aligned} & (2 \times 27.3 \mathrm{days})^{2} / r^{3}=0.00373 \\ & r=\sqrt[3]{ }\left[(2 \times 27.3 \mathrm{days})^{2} / 0.00373\right] \\ & r=90.4 \text { Earth radii }(=587600 \mathrm{~km}) \end{aligned}$ | (4) |

